

A few relativistic calculation examples.

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Example 1

Suppose we shoot a clock with a tremendous speed to Mars and back at the same speed. Mars is at that moment at **10 light minutes = 10 x 60 x 300,000 km = 180 million km** away from us (back and forth **360 million km, 20 light minutes**).

According to an observer on earth, the clock takes **40 minutes** to return. What is the state of the clock when it is set to **0:00:00** on departure?

Answer:

$$t_{clock} = \sqrt{t_o^2 - t_s^2} = \sqrt{40^2 - 20^2} = 34.64 \text{ min} = 00:34:38. \quad (\text{see } [HERE](#))$$

Example 2:

According to an observer, two missiles fly towards each other. Missile 1 flies according to the observer at a speed of **0.6 c**, missile 2 at a speed of **0.4 c**. At what speed do the missiles see each other approaching? (see [HERE](#))

For missile 1 applies:

$$v_e = \frac{c}{\sqrt{1/0,6^2 - 1}} = 0.75 c$$

For missile 2 applies:

$$v_e = \frac{c}{\sqrt{1/0,4^2 - 1}} = 0.4363 c$$

Sum of both = 1,1864 c

$$v_i = \frac{c}{\sqrt{(c/v_e)^2 + 1}} = \frac{c}{\sqrt{1/1,1864^2 + 1}} = 0.7646 c$$

So the two missiles see each other approaching with a speed of **0.7646 c**.

Example 3:

What is the kinetic energy of a particle with a speed of **0.99 c**?

$$v_e = \frac{c}{\sqrt{1/99^2 - 1}} = 7.018 c$$

$$E_k = mc \sqrt{c^2 + v_e^2} - mc^2$$

$$E_k = mc^2 \sqrt{1 + 7.018^2} - mc^2 = 6.089 mc^2$$

Example 4:

In the Large Hadron Collider (LHC) near Geneva, protons are accelerated in a ring-shaped tunnel with a circumference of **27 km** to a speed of **0.999,999,964 c**.

What is the required average field strength of the magnetic field to keep the protons in their orbit?

An electrically charged particle traverses a (part of a) circular trajectory in a magnetic field,

where the radius is determined by $r = \frac{mv_e}{Bq}$

$$\Rightarrow B = \frac{mv_e}{rq}$$

m = the mass of the proton =	$1.672,62 \cdot 10^{-27} \text{ kg}$
q = the charge of the proton =	$1.6 \cdot 10^{-19} \text{ C}$
r = the radius of the traversed circle =	$4,297.2 \text{ m}$
v_i = the speed of the proton =	$0.999,999,964 \text{ c}$

$$v_e = \frac{c}{\sqrt{(c/v_i)^2 - 1}} = 3,726.78 \text{ c}$$

$$B = \frac{1.672,62 \cdot 10^{-27} \cdot 3,726,78 \cdot 10^3 \cdot 3 \cdot 10^8}{4,297,2 \cdot 10^3 \cdot 1.6 \cdot 10^{-19}} = 2.72 \text{ T}$$

In order to keep the protons in the circular path, the LHC uses about twelve hundred 15 meter long superconducting magnets with a field strength of 8.36 Tesla.

Example 5:

How big is the centrifugal force on a proton in the Large Hadron Collider (LHC) that has to be compensated by the Lorentz force to keep the proton in its orbit. The LHC has a circumference of **27 km** and the proton has a speed of **0.999,999,964 c**.

The outward force which a revolving mass *m* exerts at a distance *r* from a center is given by:

$$F_m = \frac{mv_e^2}{r}$$

m = the mass of the proton =	$1.672,62 \cdot 10^{-27} \text{ kg}$
r = the radius of the traversed circle =	$4,297.2 \text{ m}$
v_i = the speed of the proton =	$0.999,999,964 \text{ c}$

$$v_e = \frac{c}{\sqrt{(c/v_i)^2 - 1}} = 3,726.78 \text{ c}$$

$$F_m = \frac{1.672,62 \cdot 10^{-27} \cdot (3,726.78 \text{ c})^2}{4,297.2} = 4.865,4 \cdot 10^{-7} \text{ N}$$

Example 6:

With a spaceship a clock is shot into space with a constant acceleration. An observer stays behind.

The clock is set to 00:00:00 at the moment of departure.

The acceleration is 1000 m/s^2 .

At what distance from the observer, according to that observer, does the ship reach a speed of $0,4 c$ (v_i) and what is the state of the clock at that moment?

$$v_e = \frac{c}{\sqrt{(c/v_i)^2 - 1}} = 0.4364 c$$

t_e is the state of the clock .

$$t_e = \frac{v_e}{a} = \frac{0.4364 c}{1000} = 130,931 \text{ seconds} = 36 : 22 : 11$$

$$s = \frac{v_e * t_e}{2} = \frac{a * t_e^2}{2} = \frac{1000 * 130,930^2}{2} = 8.5714 * 10^{12} \text{ m} = 8.5714 * 10^9 \text{ km} = 28,571 \text{ light seconds}$$

which is almost equal to 8 light hours .

Example 7:

What is the difference, after a hundred years, of a clock on a 100 m high tower with a clock on the ground?

Answer:

The acceleration at the surface of the earth (and at a hundred meters height) is equal to 10 m/s^2 , what is equivalent to a time gradient of (one second is equal to $3 * 10^8 \text{ m}$)
 $10 \text{ m} / (3 * 10^8 \text{ m})^2 = 1.1111 * 10^{-16} \text{ per meter}$.

100 years is equal to $100 * 365 * 24 * 3600 \text{ s} = 3.1536 * 10^9 \text{ s}$

The time difference is

$\text{height} * \text{time} - \text{gradient} * 100 \text{ years} = 100 * 1.1111 * 10^{-16} * 3.1536 * 10^9 = 35.04 * 10^{-6} \text{ s}$
so just a little more than $35 \mu\text{s}$.

Example 8:

How much does a clock in a satellite at 20,000 km altitude go faster or slower in comparison to a clock on the earth after one revolution?

Answer:

A clock at a certain height runs faster than a clock on earth.

A clock that moves with respect to a clock on earth runs slower.

The acceleration due to the attraction of a mass applies: $g(r) = \frac{GM}{r^2}$

The product of G and M for the earth is known with great accuracy.

For the earth applies: $GM = 3,986 * 10^{14} m^3 s^{-2}$

For the satellite, the centrifugal force is equal to the gravitational pull of the earth:

$$\frac{mv_e^2}{R_s} = \frac{mGM}{R_s^2} \Rightarrow v_e = \sqrt{\frac{GM}{R_s}} = \sqrt{\frac{3,986 * 10^{14}}{26,371 * 10^6}} = 3.887,8 ms^{-1}$$

The circumference of this circular orbit: $C_s = 2\pi r = 2 * 3,1416 * 26,371 * 10^6 = 165,694 * 10^6 m$

The time of a revolution: $t_{e-revolution} = \frac{165,694 * 10^6}{3.887,8} = 42.619 s = 11:50:19$

The time difference due to its speed between the clock on earth and the clock in the

satellite is: $t_i - t_e = \frac{s}{v_i} - \frac{s}{v_e} = \frac{s(v_e - v_i)}{v_i v_e} = \frac{s(1 - v_i/v_e)}{v_i}$

$$v_e = \frac{v_i}{\sqrt{1 - v_i^2/c^2}} \Rightarrow t_i - t_e = \frac{s}{v_i} (1 - \sqrt{1 - v_i^2/c^2}) \approx \frac{s v_i}{2c^2}$$

The time difference due to its speed after one revolution:

$$t_i - t_e = \frac{165,694 * 10^6 * 3.887,8}{2(3 * 10^8)^2} = 3,5788 * 10^{-6} s$$

If you do not like approaches, you can calculate the result using the Ivy calculator from

Google: $t_i - t_e = \frac{s}{v_i} (1 - \sqrt{1 - \frac{v_i^2}{c^2}}) = 3,5788 * 10^{-6} s$

In **Example 7**, g is approximately a constant. However, that does not apply here. We therefore integrate $g(r)$ from R_0 to R_s .

$$g[R_0 - R_s] = \int_{R_0}^{R_s} \frac{GM}{r^2} dr = \left[\frac{GM}{r} \right]_{R_0}^{R_s} = GM \left(-\frac{1}{R_s} + \frac{1}{R_0} \right) = GM \left(\frac{R_0 - R_s}{R_s R_0} \right) m^2 s^{-2}$$

$$g[R_0 - R_s] = 3.986 * 10^{14} \left(\frac{6.371 * 10^6 - 26.371 * 10^6}{6.371 * 10^6 * 26.371 * 10^6} \right) = -47.450 * 10^6 m^2 s^{-2}$$

The time difference of the two clocks due to gravity after one revolution is:

$$\frac{-47,45 * 10^6}{(3 * 10^8)^2} * 42.619 = -22,47 * 10^{-6} s$$

The total time difference after one revolution is therefore

$$3,5788 * 10^{-6} - 22,47 * 10^{-6} = -18,89 * 10^{-6} s \text{ , so almost } -19 \mu s$$

To be continued.

In this article/discussion I will show that it is better for a good understanding of the theory of relativity to take more distance from observation to interpret the observation data. In addition, the theory of relativity can be better understood.

Disclaimer

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